

Focus errors from tracking sodium layer altitude variations with laser guide star adaptive optics for the Thirty Meter Telescope

CfAO Fall Retreat, Nov. 2006, Yosemite CA

Glen Herriot^a, P. Hickson^b, S. Davis^b, R. Clare^f,

B. L. Ellerbroek^c, J.-P. Véran^a, C. Y. She^d, D. Looze^e

^aNational Research Council Canada – Herzberg Institute of Astrophysics,

^bUBC Canada, ^cTMT, ^dColorado S.U., ^eU. Mass. ^fKeck Obs.

Motivation & Outline.

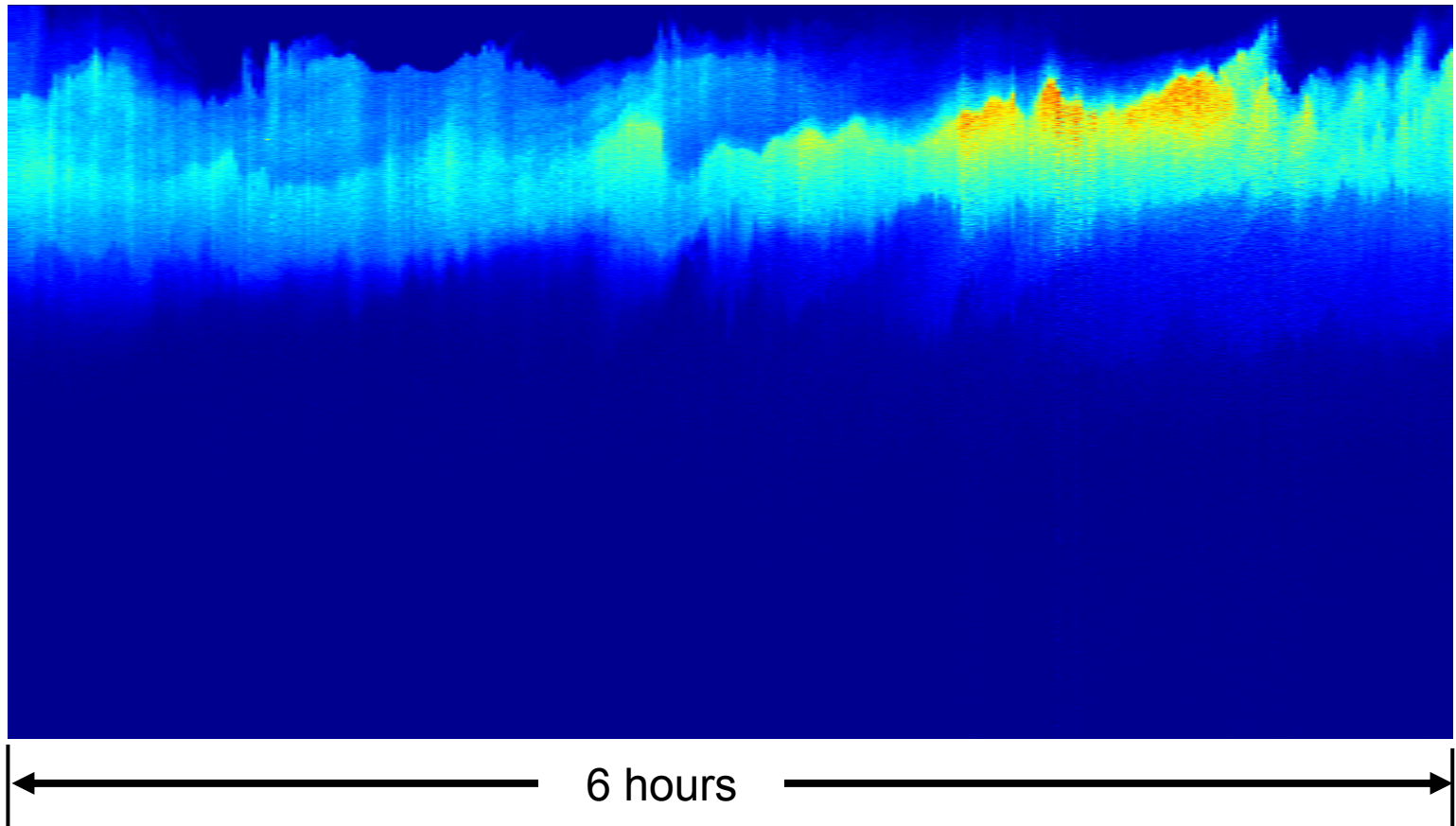
- ◆ Sodium altitude variation introduces defocus errors for LGS AO on TMT and other large aperture telescopes
- ◆ Sensitivity grows $\sim D^2$
- ◆ Focus sensing via natural guide stars, with good sky coverage, leaves residual servo errors
- ◆ Residual WFE = $f(\text{Natural Guide Star sky coverage})$
- ◆ Na Layer thickness reduces signal vs noise on LGS WFS via spot elongation
- ◆ Pulsed lasers and novel CCD clocking scheme help both problems

Lidar measurements of Sodium Layer

100 km

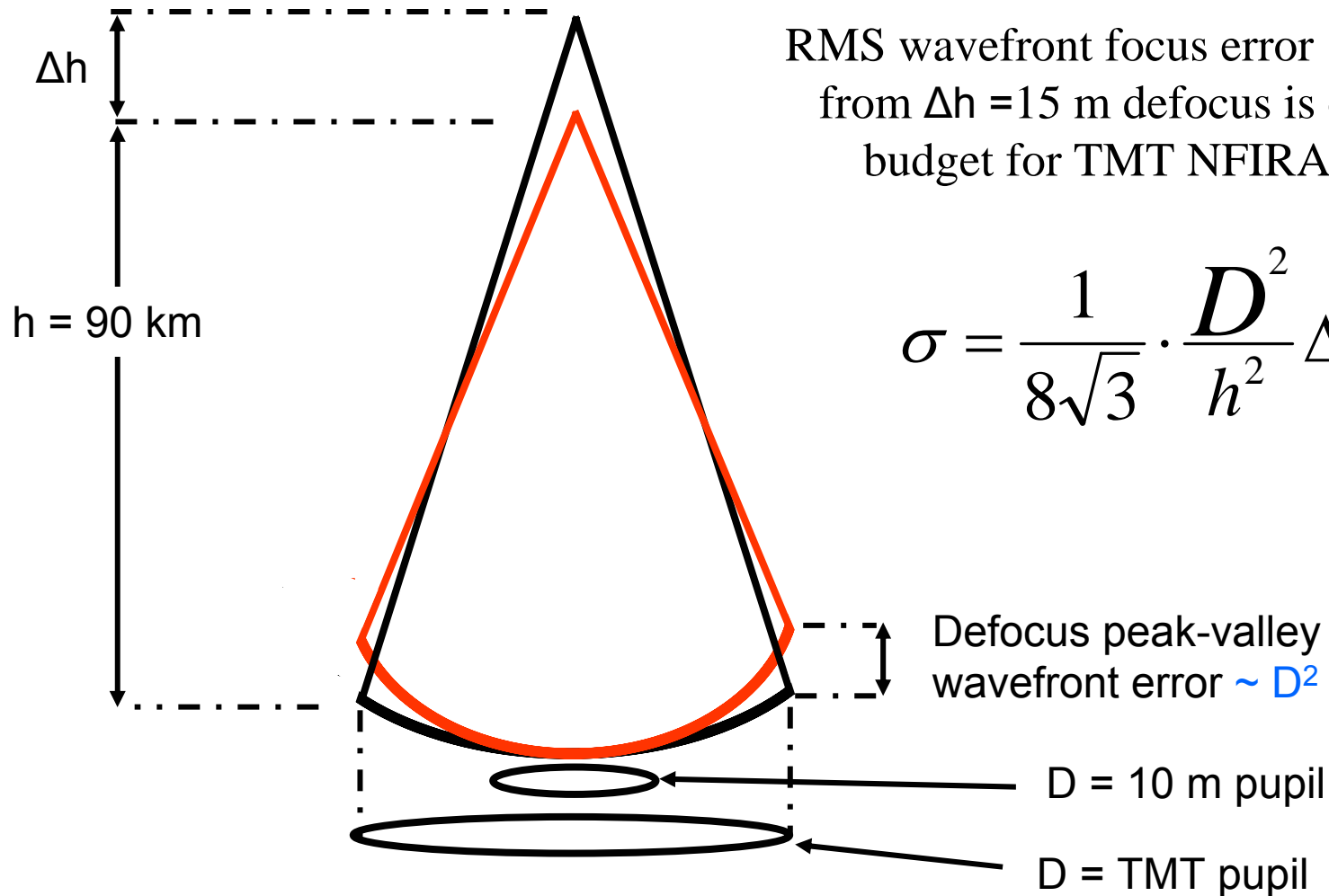
80 km

60 km



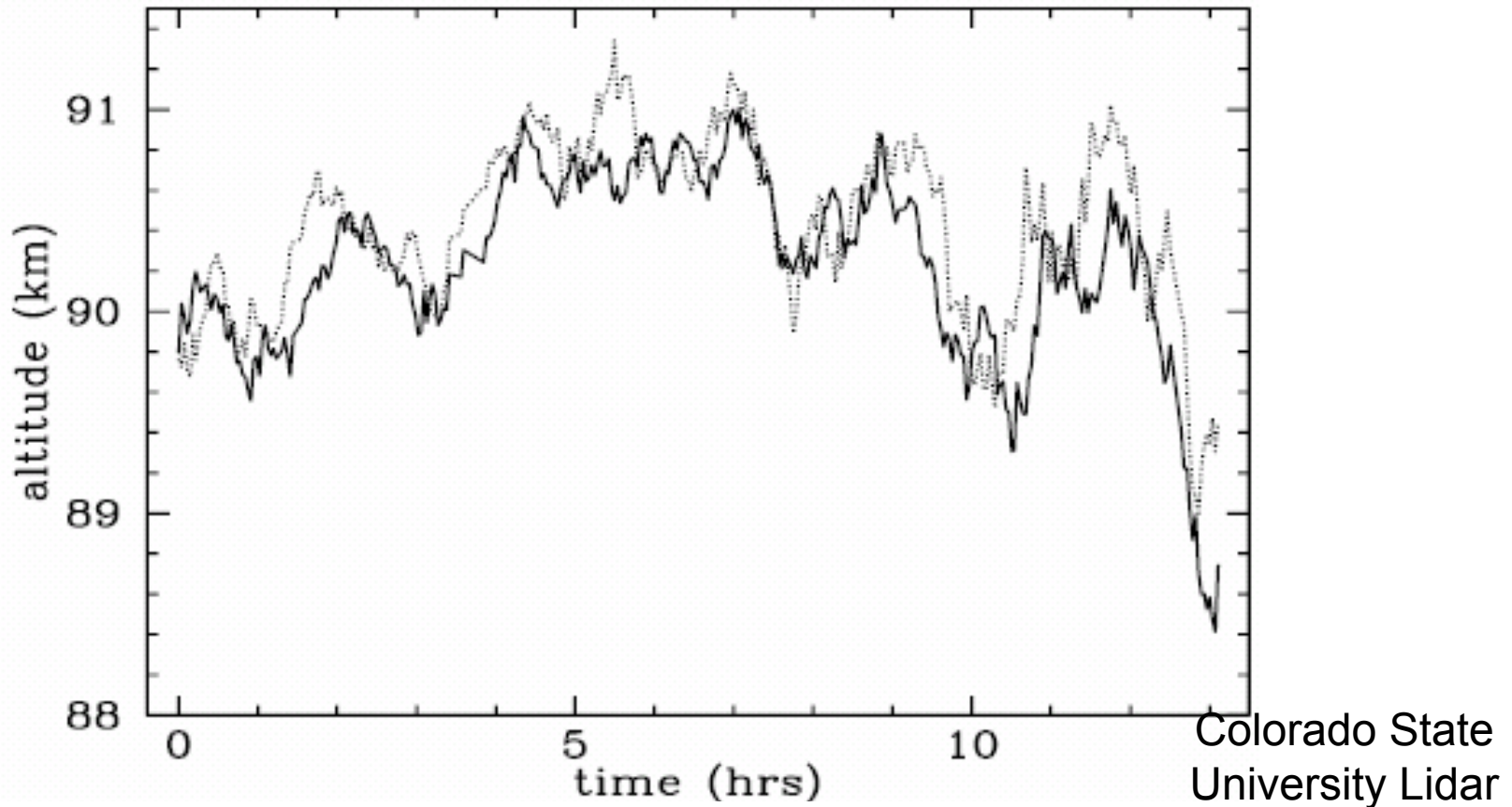
◆ Purple Crow Lidar, University of Western Ontario

Defocus Wavefront error $\sim D^2$



Sodium altitude time series

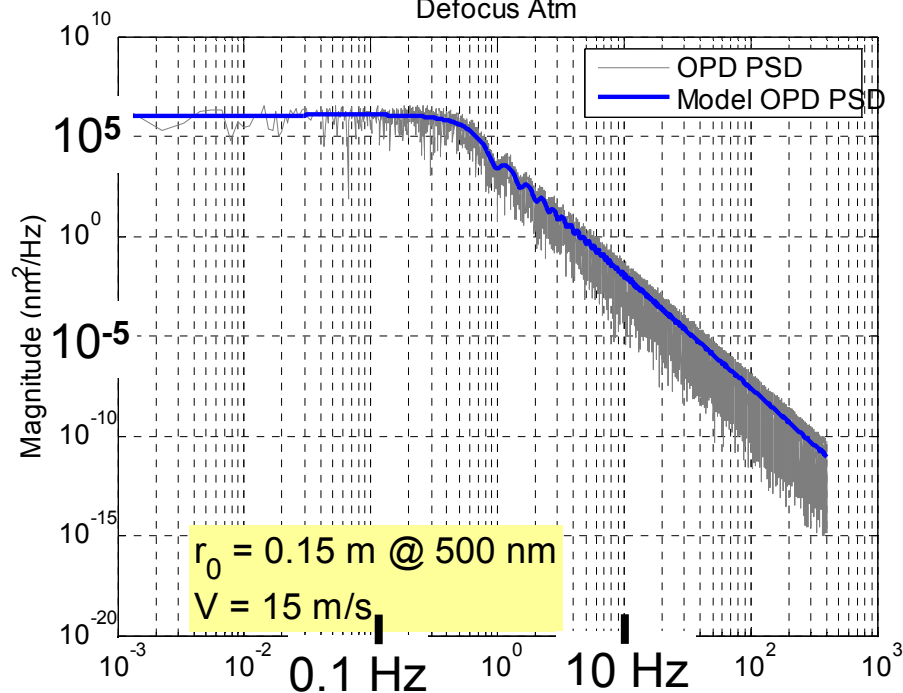
- One night, at zenith and 30 degrees off-zenith



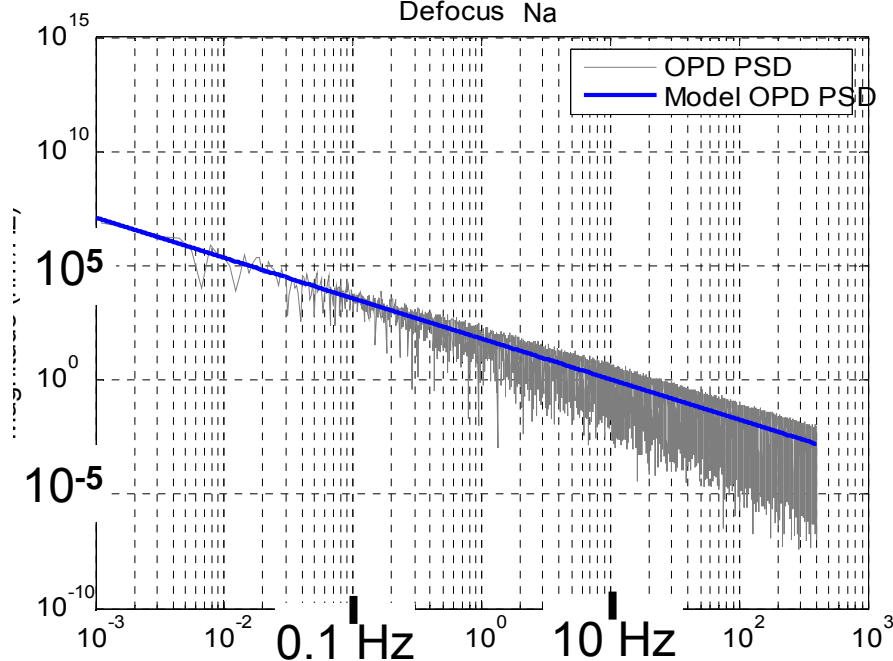
Temporal frequencies overlap: turbulence focus and Na altitude defocus

Atmospheric Focus
PSD nm^2/Hz

Defocus Atm



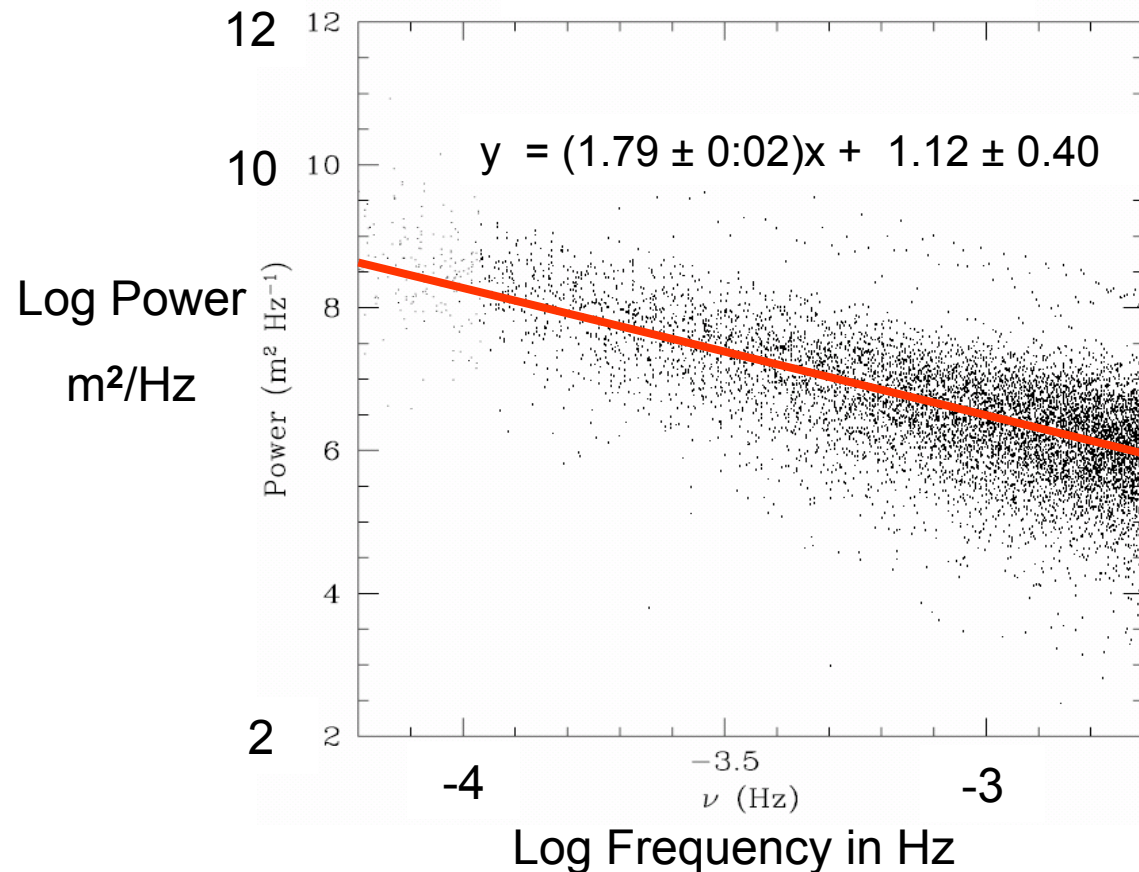
Defocus Na



Focus PSD nm^2/Hz from Na altitude
variations at 30-m telescope

Power Spectrum of Na Altitude

28 Nights in 2002 with
>7 continuous hours



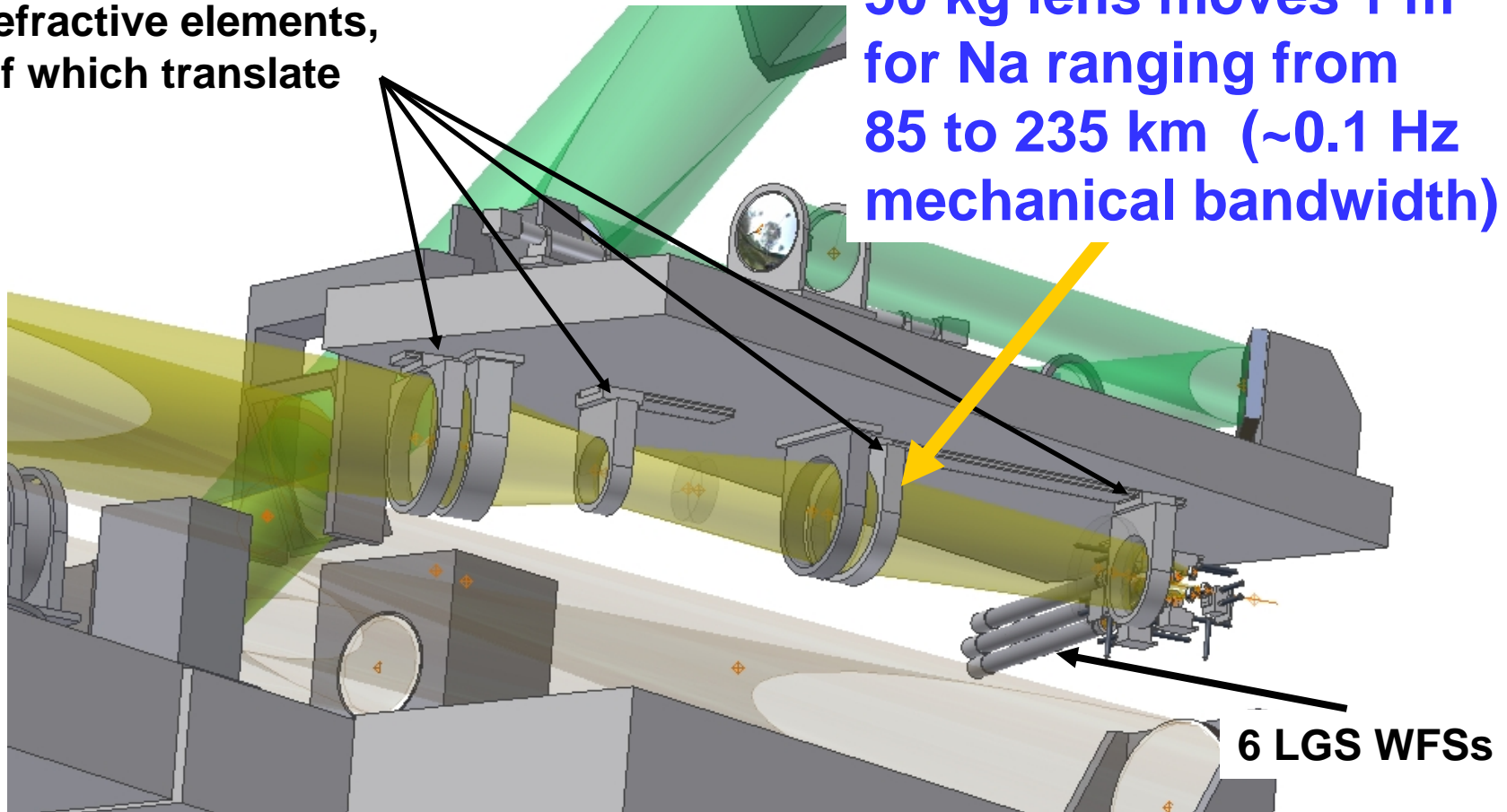
- ◆ Only longer data series (>7 hrs) included
- ◆ Long time → Narrower windowing function in Fourier domain
- ◆ ∴ Steeper PSD slope than we showed in Orlando.
- ◆ ∴ More benign high frequencies

Colorado State
University Lidar

TMT NFIRAOS' Laser WFS Components

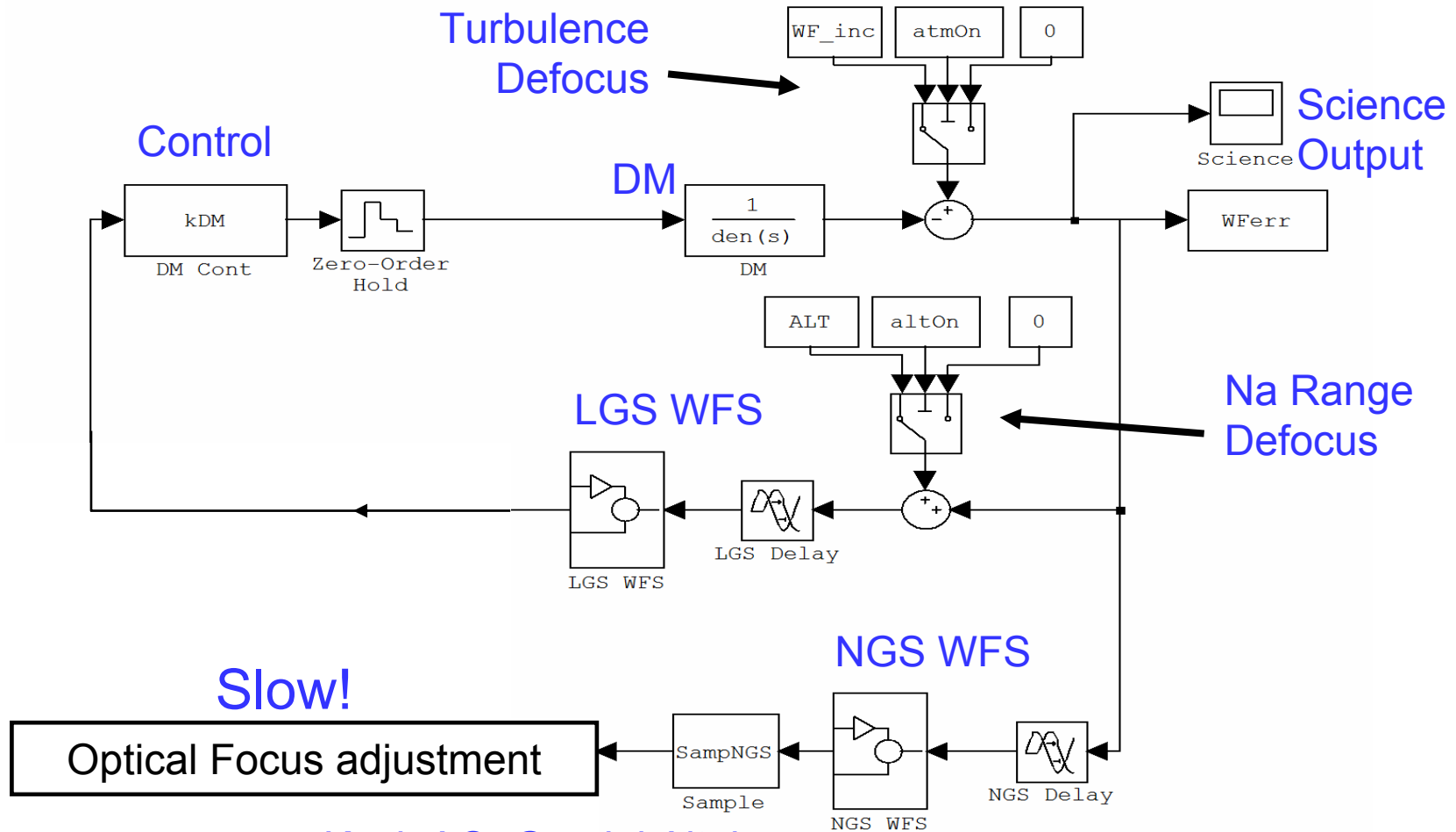
6 refractive elements,
4 of which translate

50 kg lens moves 1 m
for Na ranging from
85 to 235 km (~0.1 Hz
mechanical bandwidth)



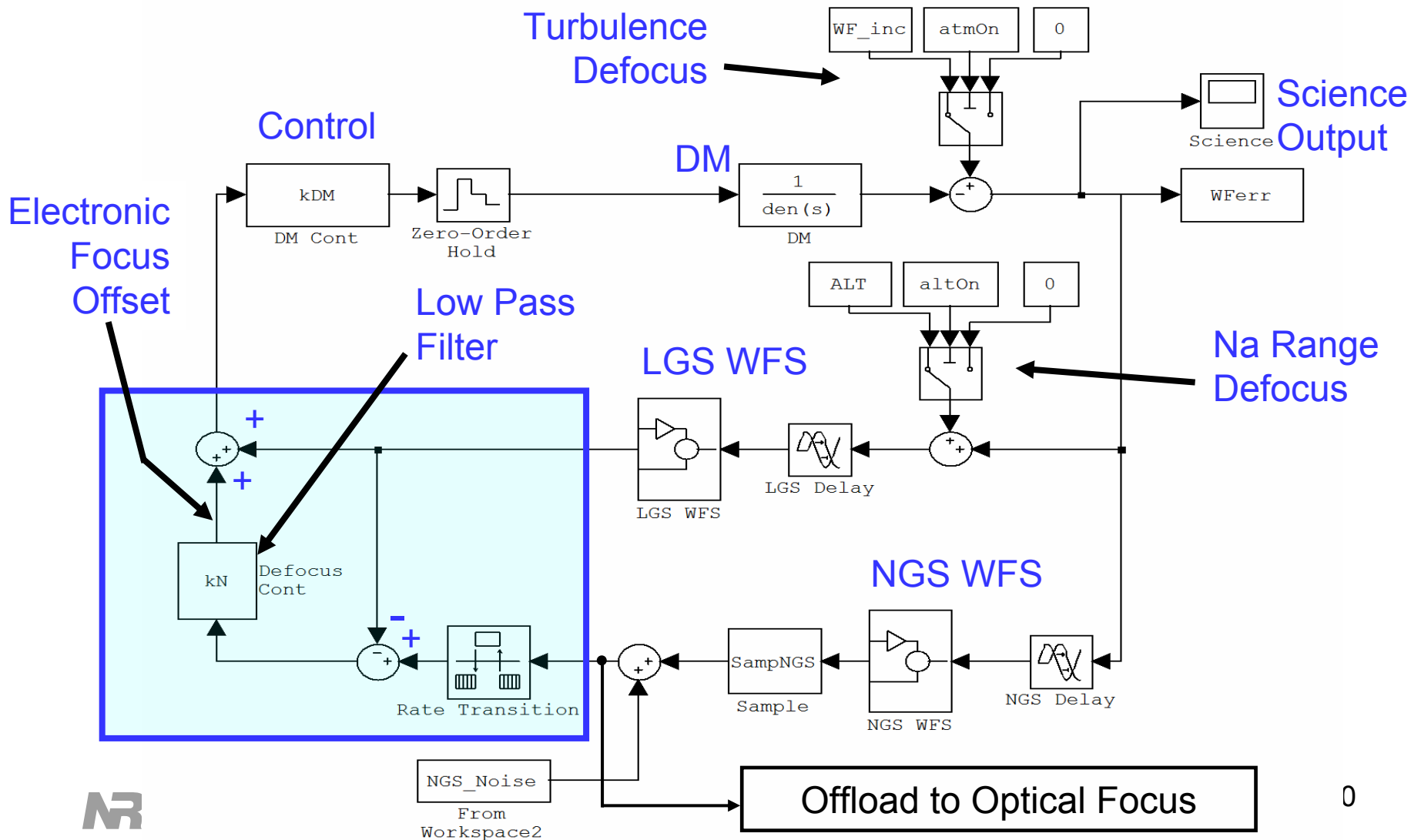
Simulink Model

8-m to 10-m Telescope Na Tracking

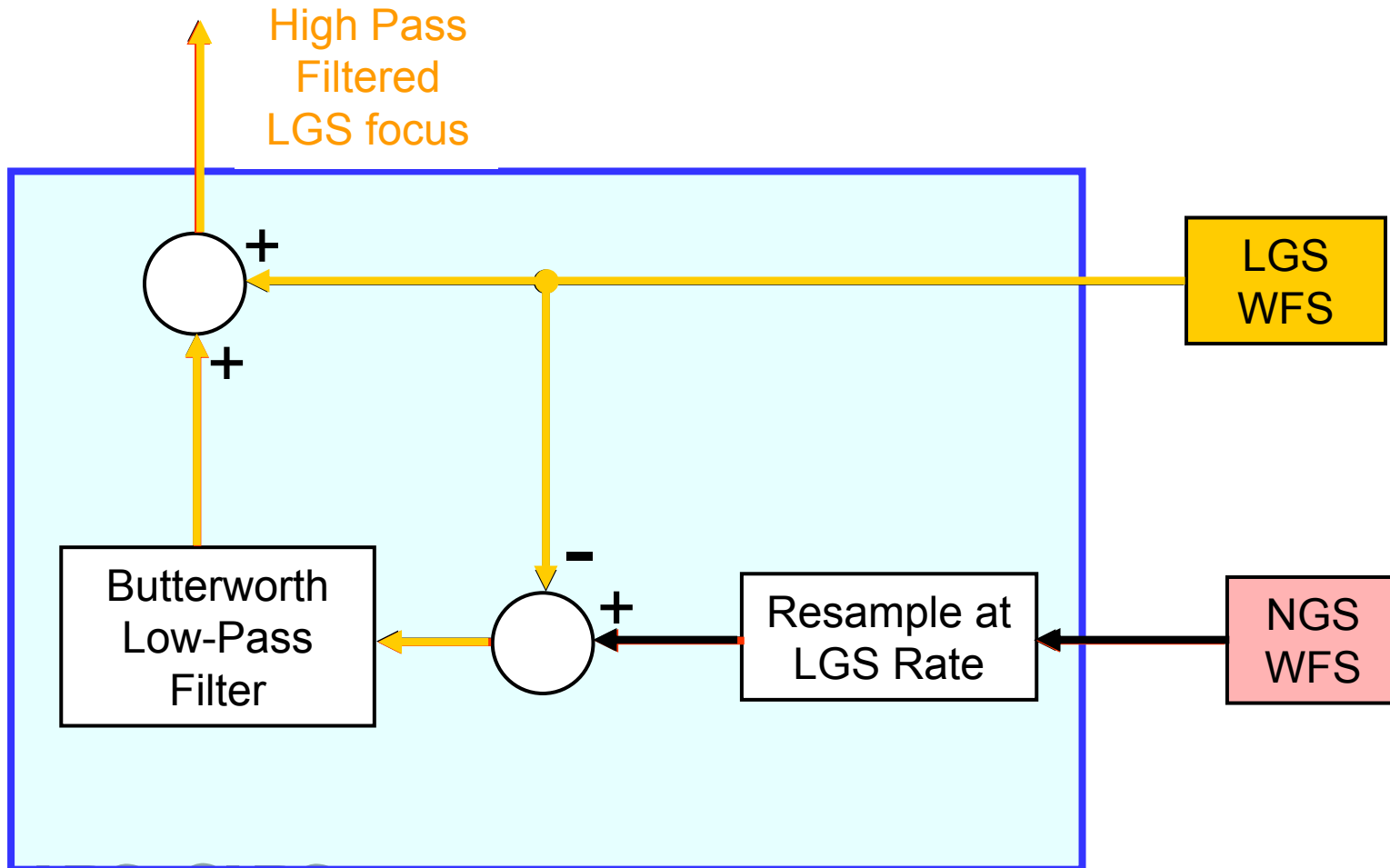


e.g. Keck AO, Gemini Altair,
Gemini South MCAO

Sodium Layer Electronic Focus Offset for 30-m Telescope

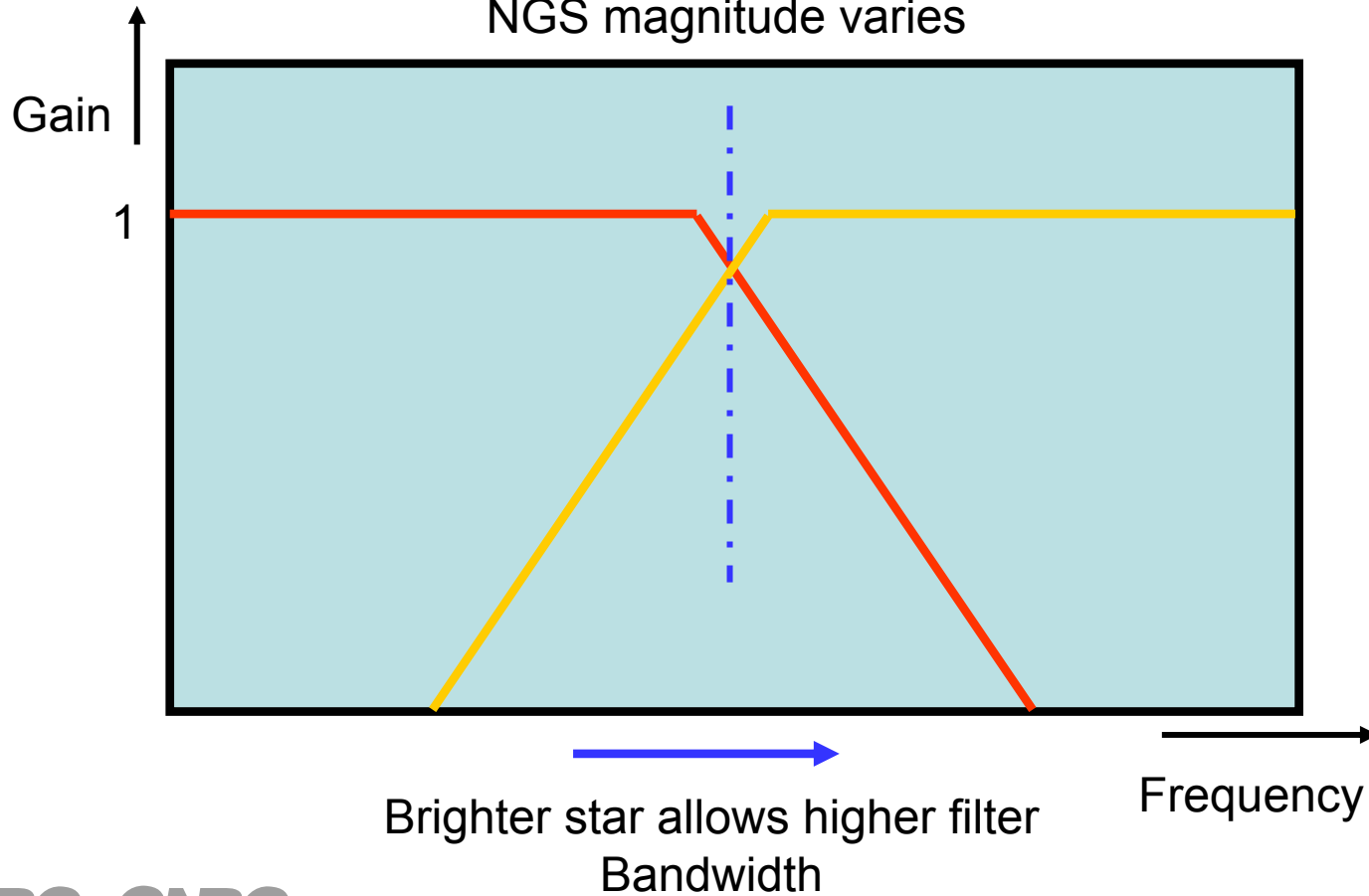


Electronic Focus Offset Detail

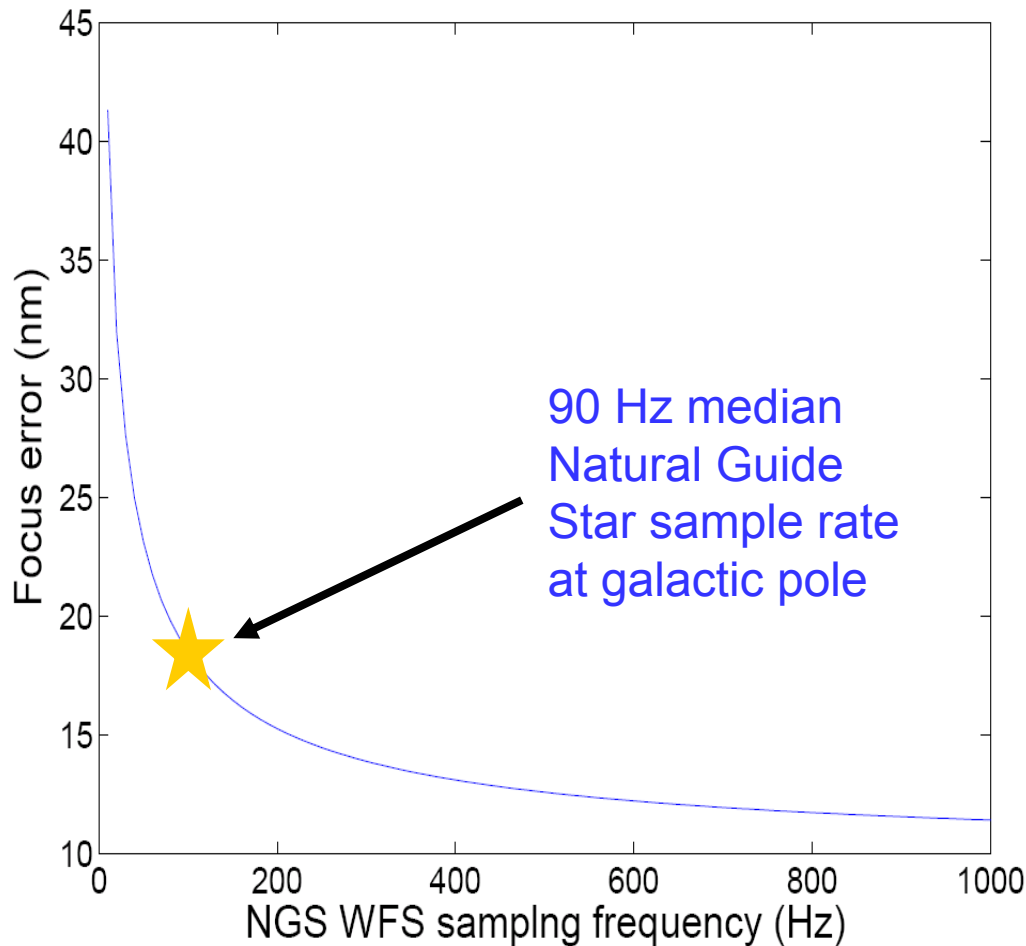


Why such an arcane scheme?

- ◆ NGS and LGS filters are exactly complementary
- ◆ Single parameter, Filter BW, adjusts blending as NGS magnitude varies

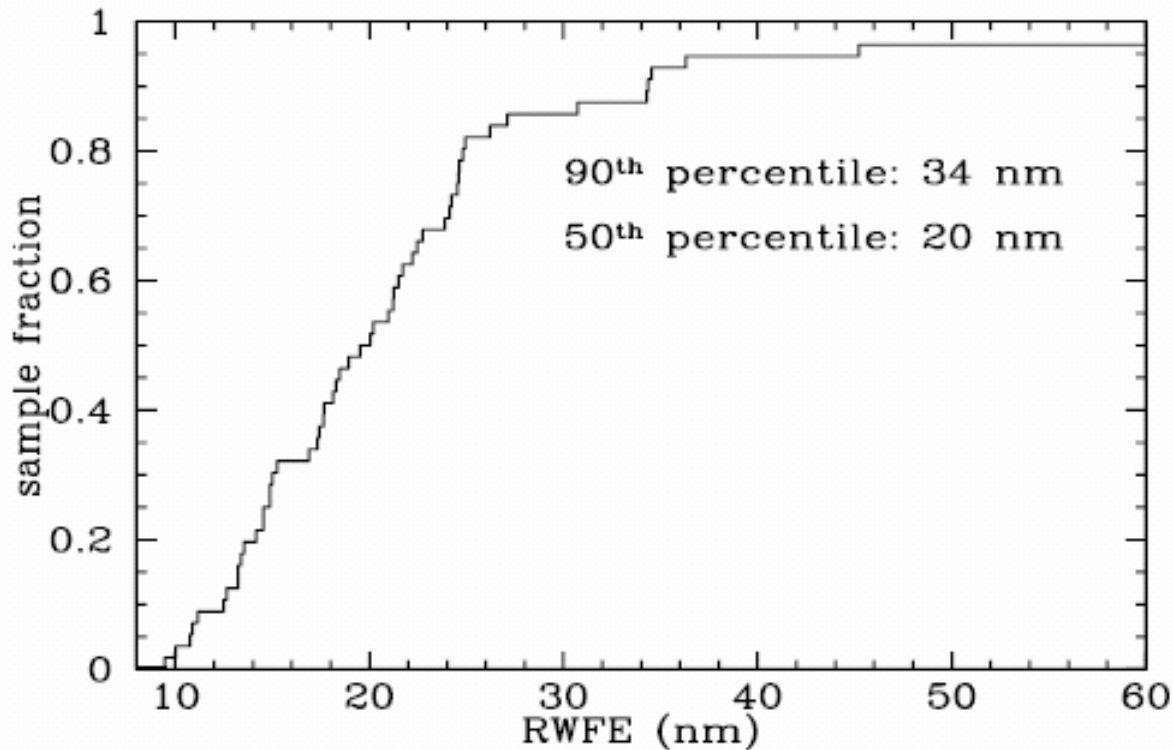


Na Defocus vs sample rate using WFS electronic focus offsets



- ◆ Based on Tip-Tilt error budget & Sky coverage models for TMT NFIRAOS
- ◆ Optimization of frame rate to balance servo lag, read and photon noise, measurement errors, anisoplanatism, and sodium tracking error
- ◆ R. Clare, TMT, SPIE Orlando, 6272-107

Residual error after electronically refocusing via NGS measurements



Latest results from the CSU Lidar

56 time series, each >7 hours duration

90 Hz electronic focus offsets applied to LGS WFSs

50% sky coverage at galactic pole.

Sodium layer altitude tracking

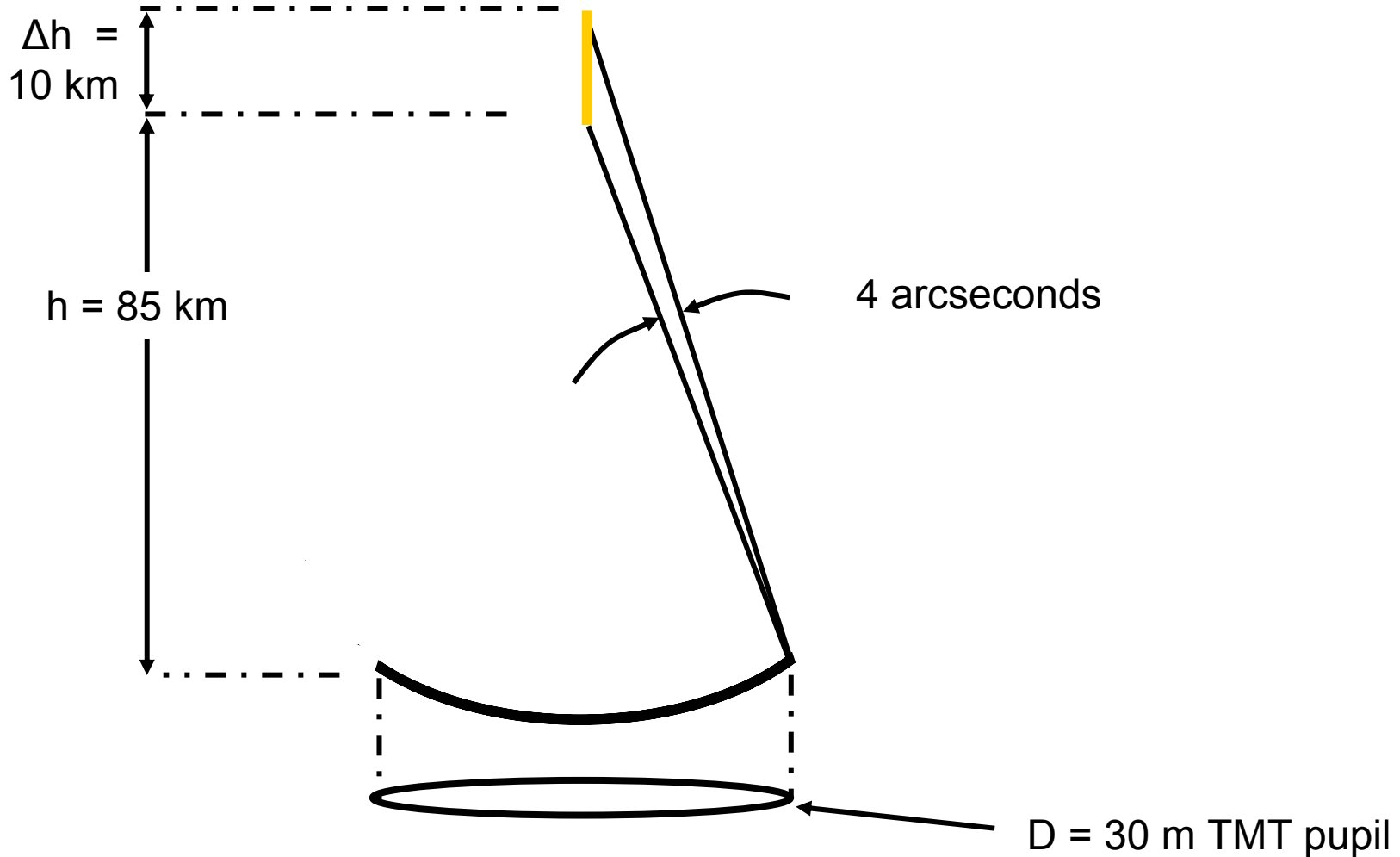
- LGS WFS confounds Na layer altitude change with turbulence
- Need to rely on moderate bandwidth NGS WFS for focus tracking
- On $D=30\text{m}$, 1 m of tracking error = 8 nm rms (D^2 effect)
- Lidar data available with temporal resolution of 2 min
- From each measurement, a PSD is derived and extrapolated
- Rejection transfer function is applied (90 Hz frame rate, electronic compensation)
- Residual is computed
- Caveat: 3 decades of temporal frequency extrapolation
 - Higher time-resolution data planned for further studies



TMT

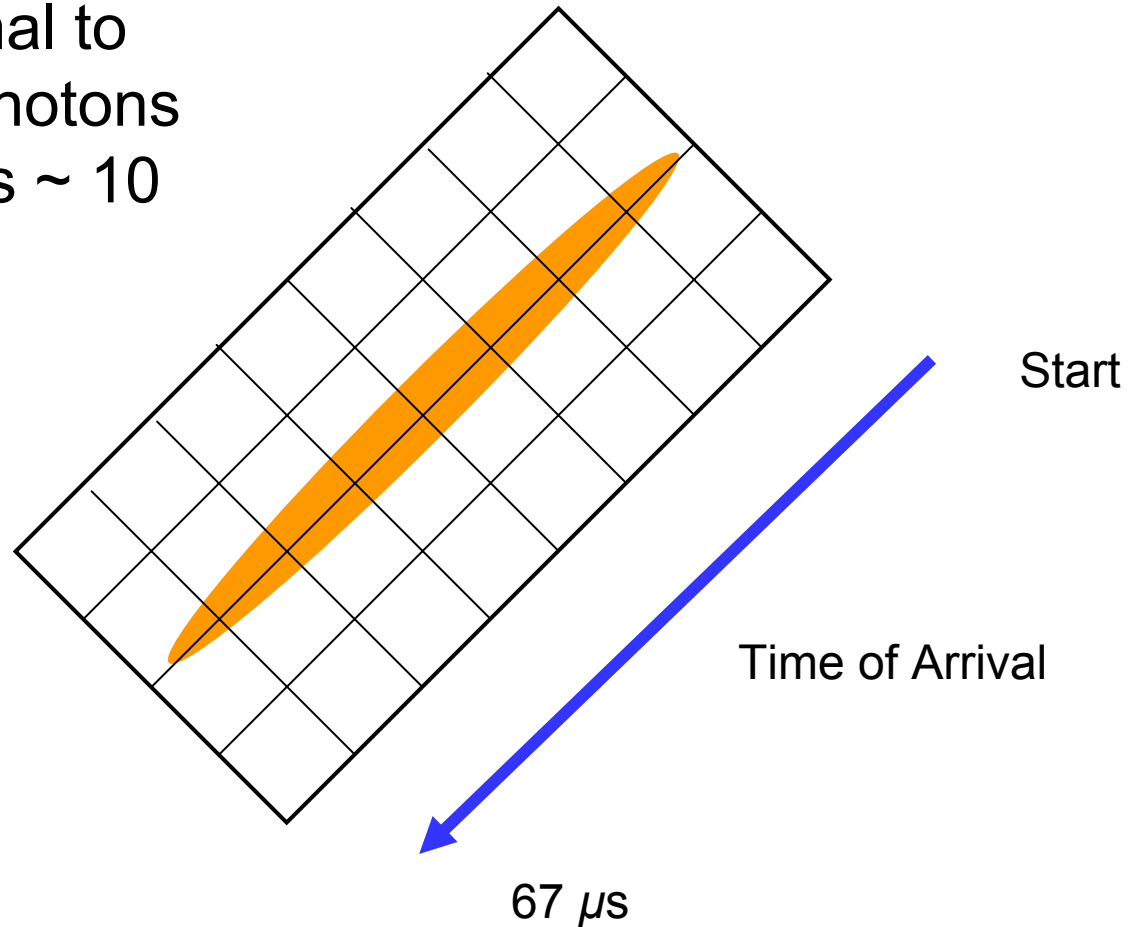
THIRTY METER TELESCOPE

WFS Spot elongation 4 arcseconds

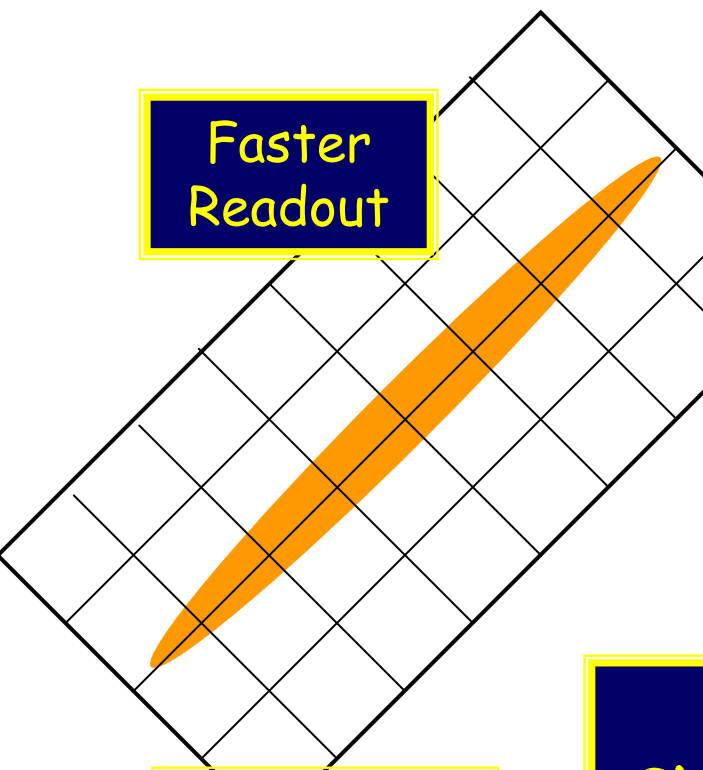


Elongated Pattern on CCD

- Reduced signal to noise since photons spread across ~ 10 pixels.

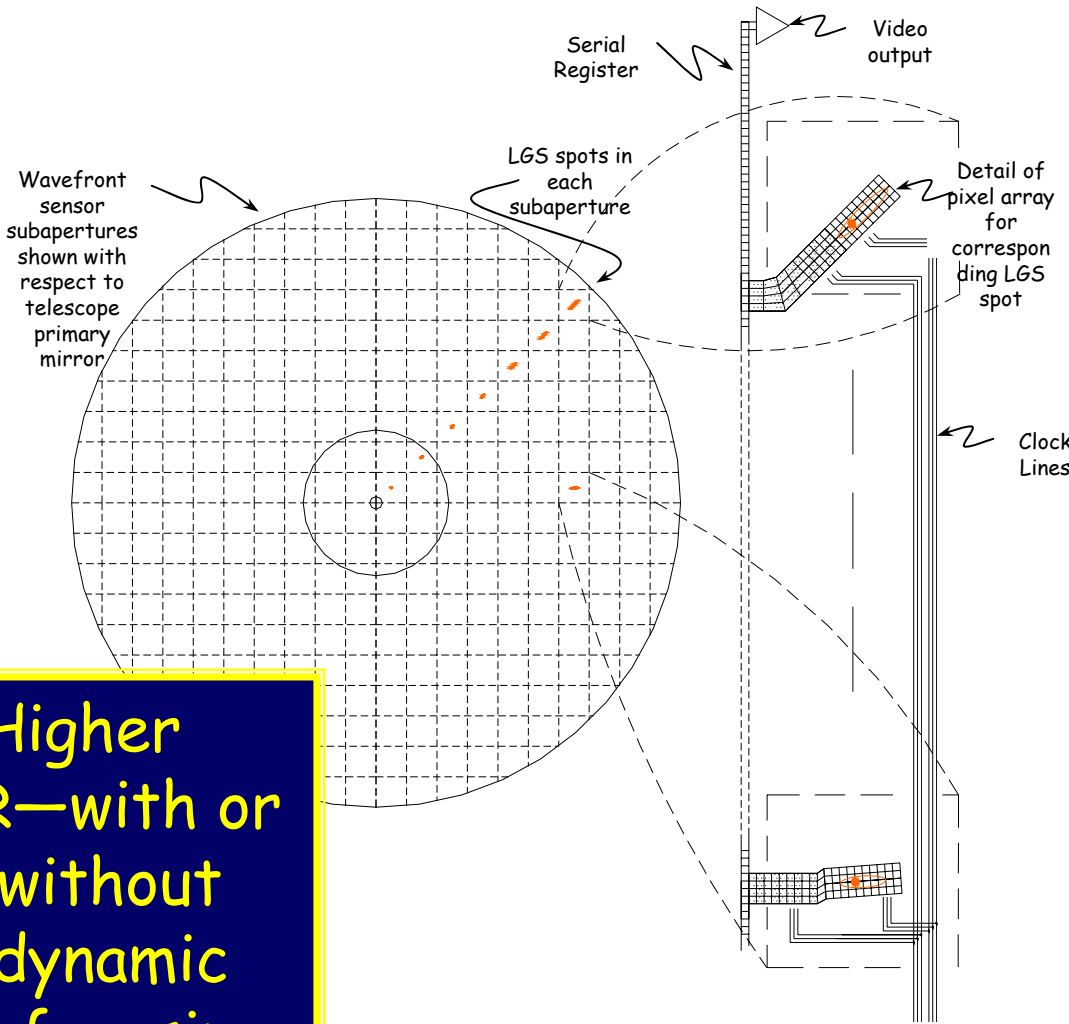


Polar coordinate CCD Pixel Geometry for Laser Guidestars



**Faster
Readout**

**Lower
Readout
Noise**

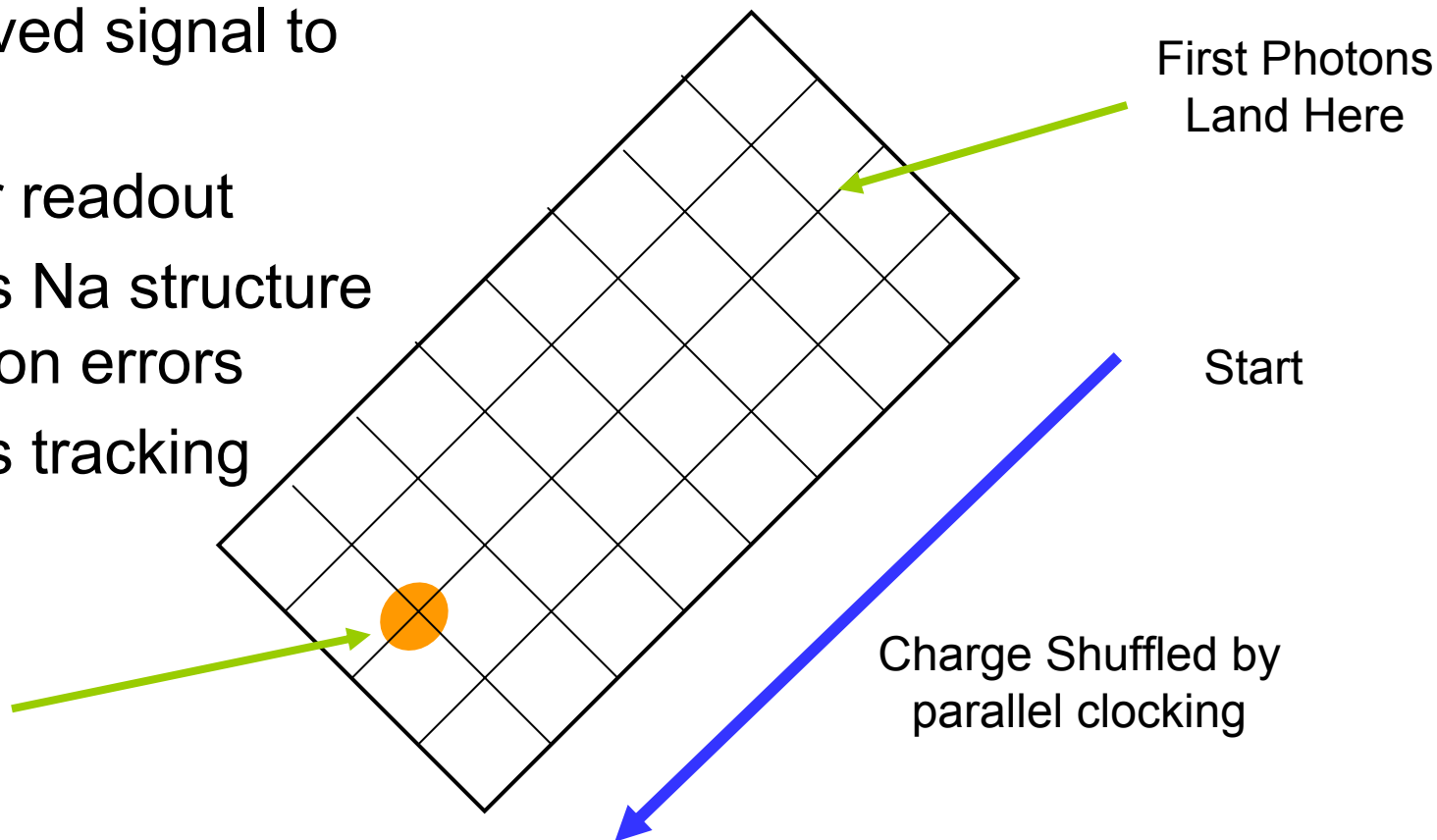


**Higher
SNR—with or
without
dynamic
refocusing**

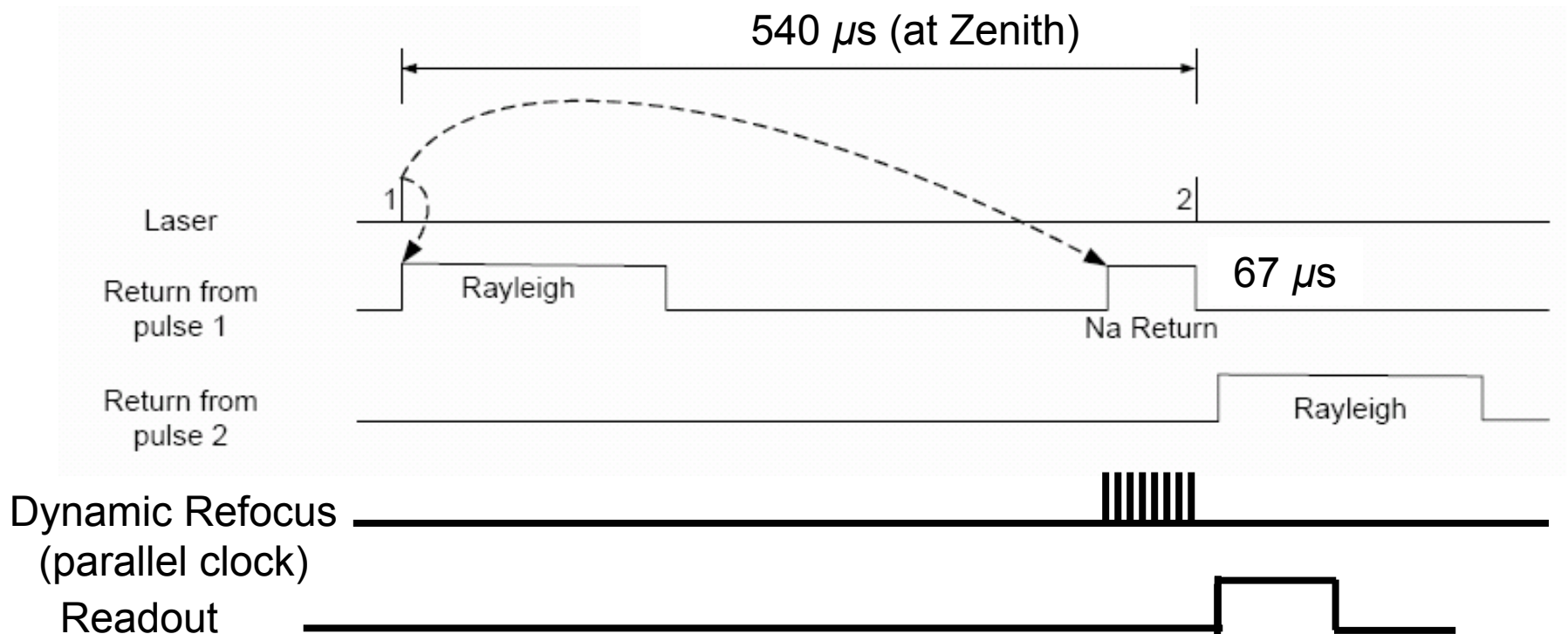
Dynamic Refocusing on CCD

- ◆ Improved signal to noise
- ◆ Faster readout
- ◆ Avoids Na structure variation errors
- ◆ Avoids tracking errors

Last photons land here 68 μ s later, and add to the earlier ones

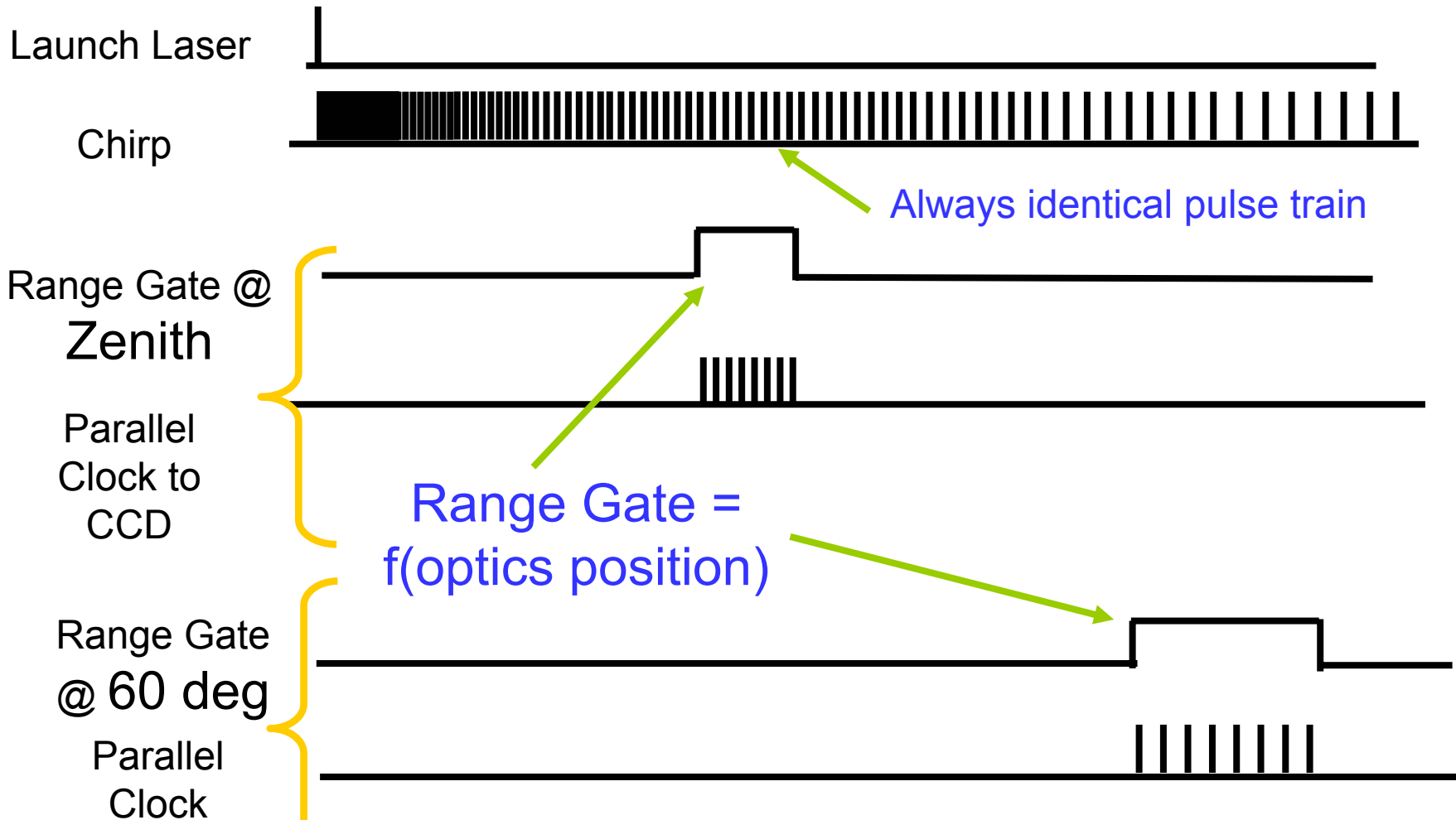


Pulsed Laser Timing at Zenith



- Sean Adkins

Chirped Clock



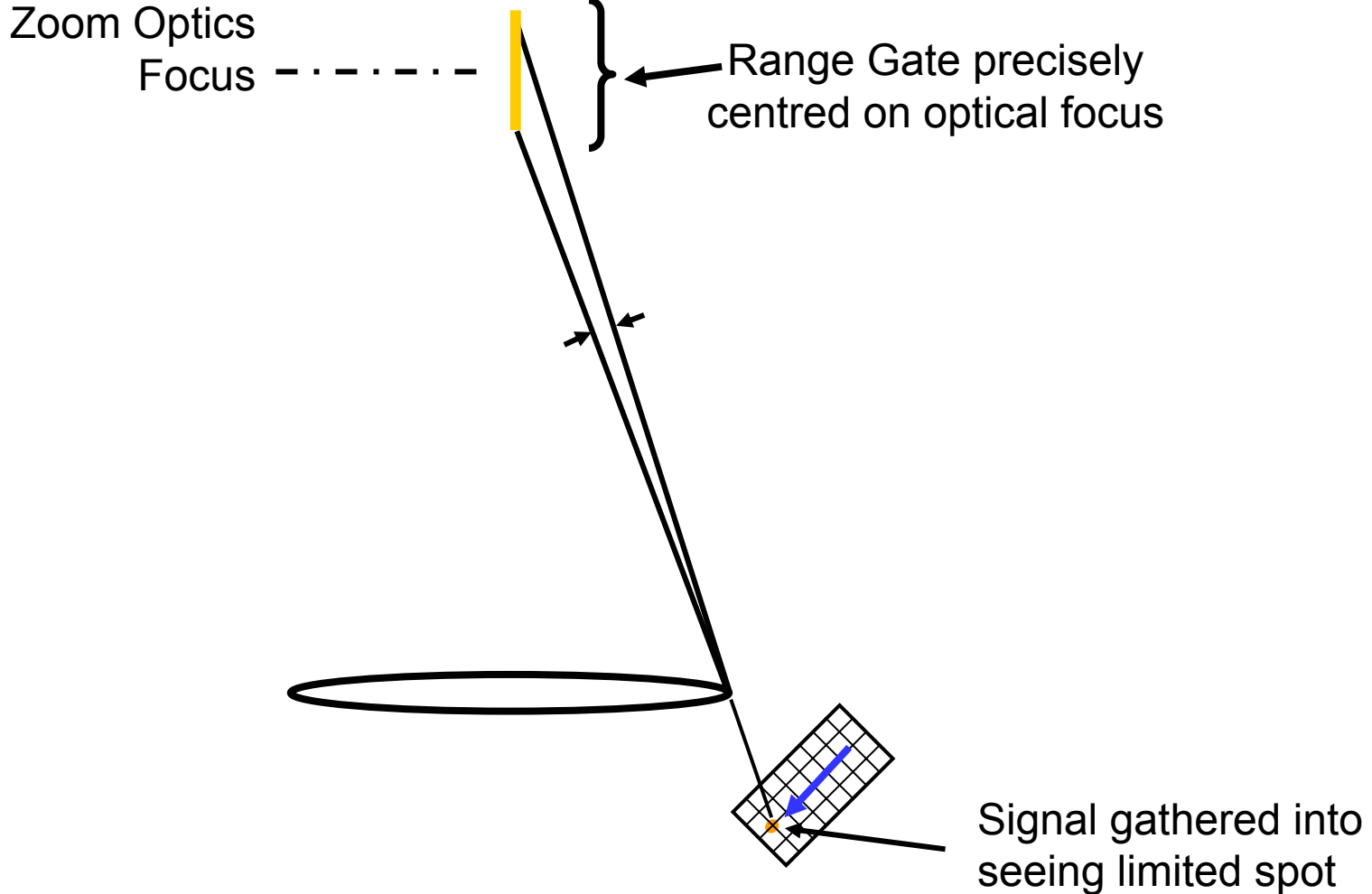
Why does chirp fix focus tracking error?

- ◆ Return from a given range, arrives at a known time and is collected correctly on CCD.
- ◆ Range Gating of chirp pulsed train onto CCD at time determined by current setting of zoom optics.
- ◆ Good calibration of mechanical position to range gate interval needed.
- ◆ Does not depend on accurate mechanical focusing of zoom optics to track sodium altitude.
- ◆ If zoom is severely wrong, flux lost, but no focus error.

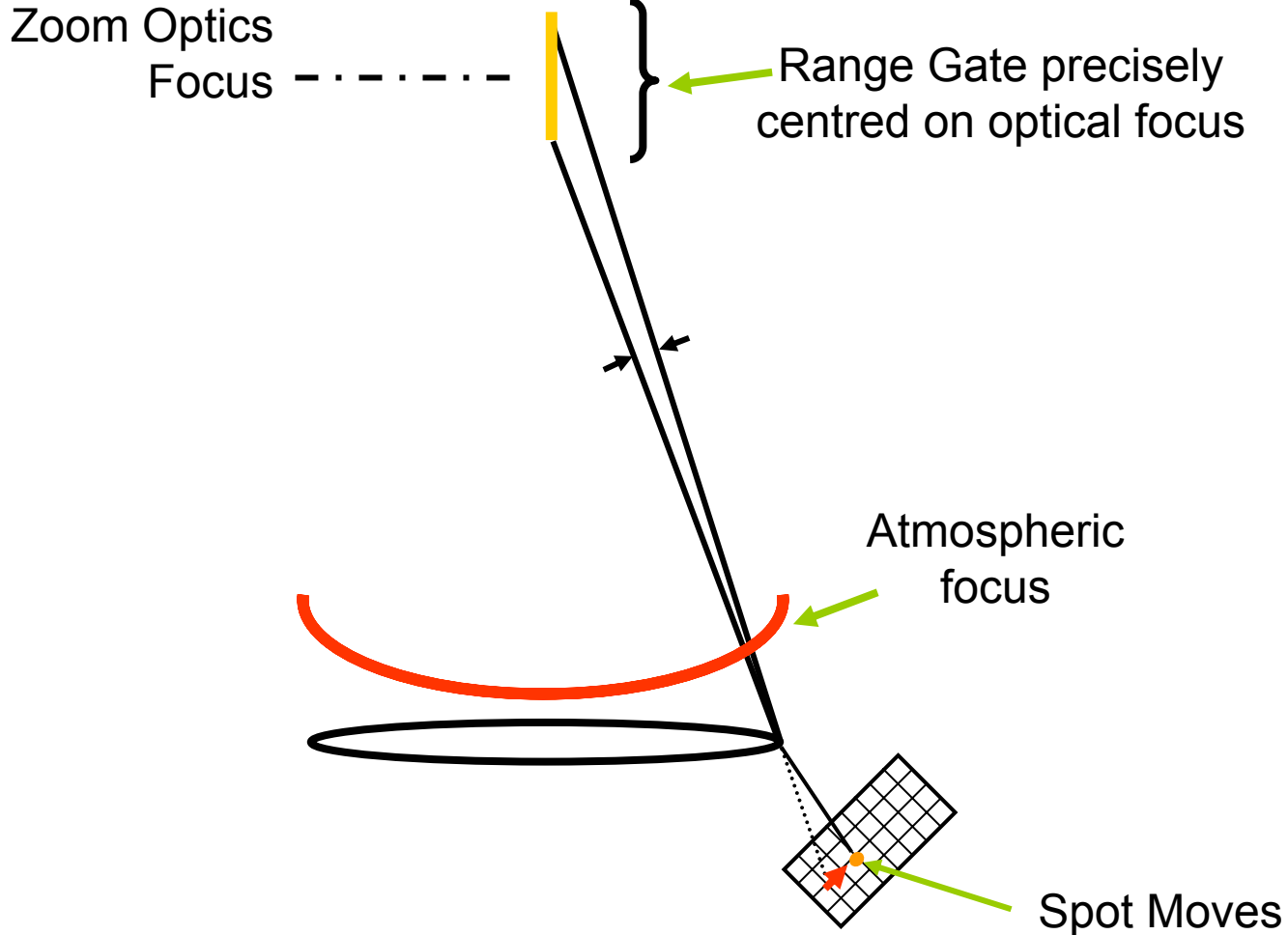


TMT Range Gated Chirp -- ideal situation

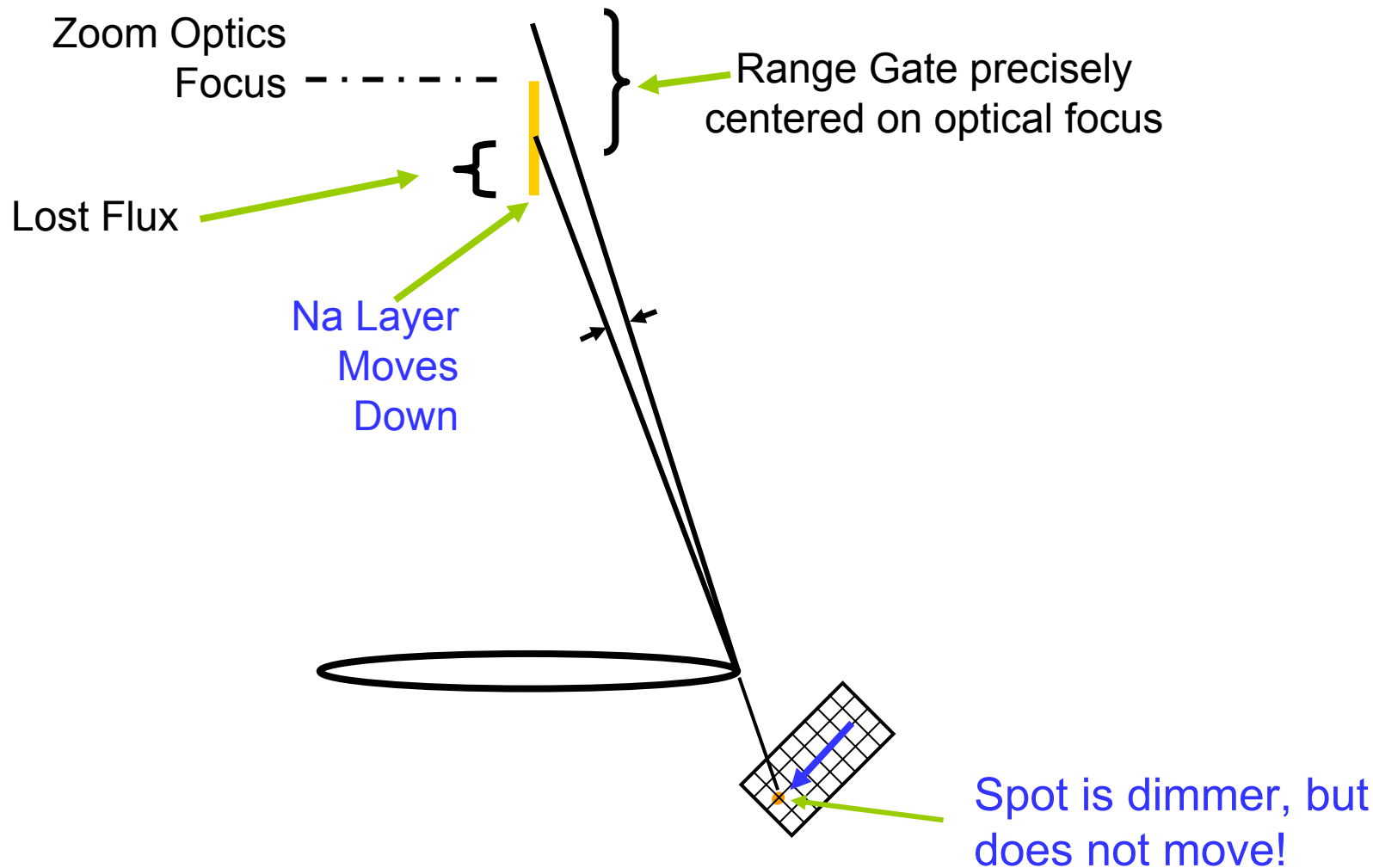
THIRTY METER TELESCOPE



Range Gated Chirp -- Turbulence



Chirp: Sodium Layer moves



Summary

- ◆ Sodium Layer mean altitude and structure varies on short time scales, causing measurement errors.
- ◆ Spot elongation reduces signal to noise ratio.
- ◆ Range-gated chirp-clocked polar-coordinate CCD mitigates both these problems.
 - But depends on a pulsed ($\sim 3 \mu\text{s}$) laser